

Solution of Exercises

Lecture Optical design with Zemax– Part 3

3	Aberrations.....	1
3.1	Performance of an achromate	1
3.2	Aplanatic lens.....	4
3.3	Transverse aberration curve.....	7
3.4	Aspherical lens.....	9

3 Aberrations

3.1 Performance of an achromate

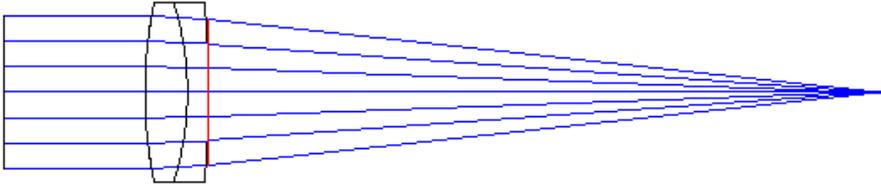
Load a classical achromate out of a vendor catalog with focal length $f = 100$ mm.

- a) What is the numerical aperture of the system in the image side ? Is the system diffraction limited ?
- b) Determine the longitudinal aberrations for the colors eFC'. What is the largest spherical aberration in the green ? What is the height of the corresponding ray in the pupil ? What is the largest difference of the axial color aberration ?
- c) Calculate the Seidel surface contributions of the system in the desired orientation and for the reversed lens.
- d) Determine the range of finite field angles, for which the original achromate is diffraction limited, if $\lambda = 546$ nm and a reduced aperture diameter of 15 mm is considered.

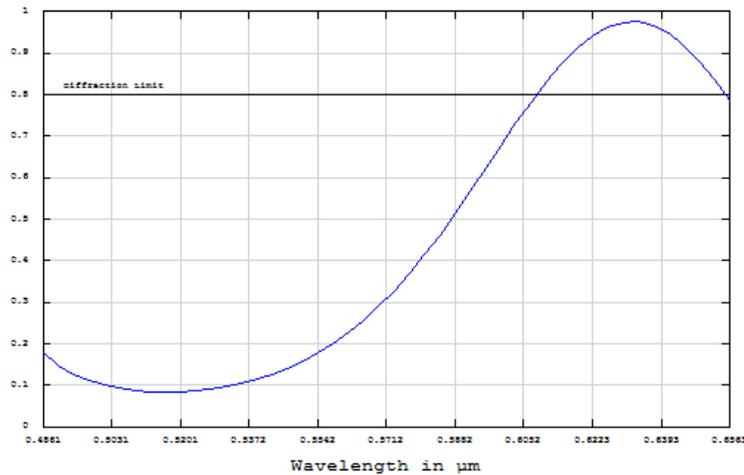
Solution:

- a) Setup (Melles Griot, AAP-100.0 - 25.4)

Lens Data Editor						
Edit Solves View Help						
Surf	Type	Comment	Radius	Thickness	Glass	Semi-Diameter
OBJ	Standard		Infinity	Infinity		0.0000000
*	Standard	AAP-100.0-25.4	62.0050000	6.0000000	N-BK7	12.7000000 U
2*	Standard		-42.379000	3.0000000	SF2	12.7000000 U
3*	Standard		-133.28400	95.6878729	M	12.7000000 U
IMA	Standard		Infinity	-		4.487E-003



The numerical aperture is $NA = 0.107$. The system is diffraction limited for wavelength $\lambda > 608$ nm (rms vs wavelength, Strehl criterion, 9 rays, 100 wavelengths):

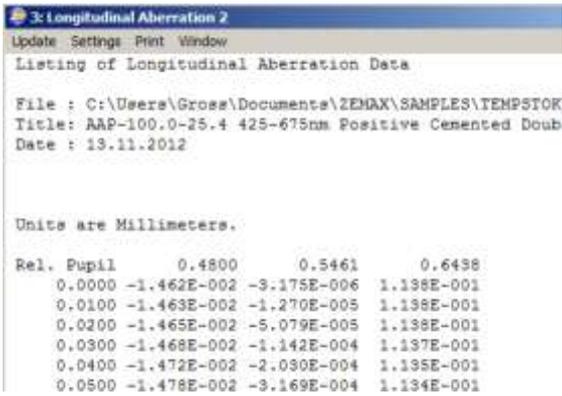


b) Take eF'C' as wavelengths and calculate the longitudinal aberration curves.



The largest spherical aberration in the green can be obtained by the text output and is $s' = -0.0377$ mm. this value corresponds to a relative pupil height of 0.76.

From the first row of the text output we get the maximum difference along the axis between blue and red of $128.4 \mu\text{m}$.

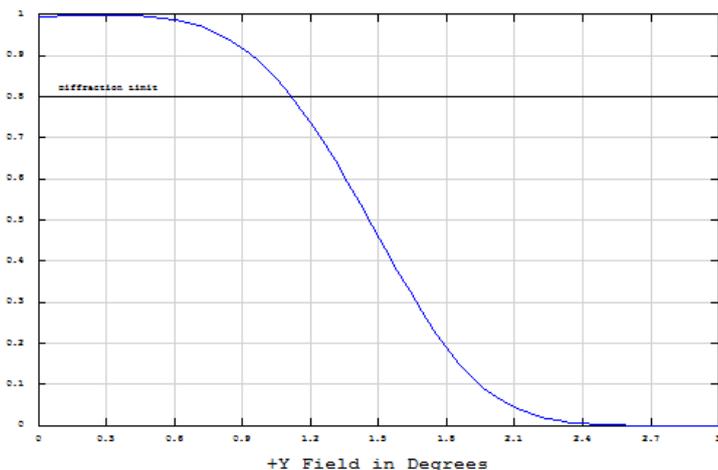


c) Seidel aberration bar chart for the original and the reversed system:

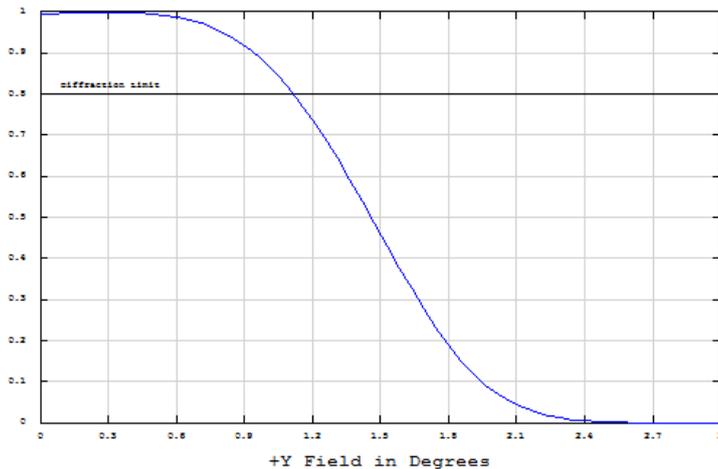


It is seen, that the the system is quite bad for the reversed lens. Therefore the catalog component should be used only with incoming collimated light.

d) With the rms-menu vs. field size and the Strehl ratio criterion with more rays, we get the following drawing after a quick focus for the axis point only:



It can be seen, that the diffraction limit is violated for field angles $w > 1.1^\circ$.



It can be seen, that the diffraction limit is violated for field angles $w > 1.1^\circ$.

3.2 Aplanatic lens

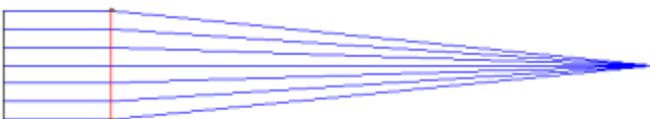
Consider a collimated incoming beam with wavelength 500 nm and diameter 10 mm. This bundle should be focussed by a perfect lens of focal length $f = 50$ mm.

- Place an aplanatic-concentric lens shortly behind the ideal lens with the material SF57. What is the resulting numerical aperture in the image space? Show at least two different methods to find the best image position.
- Show that the spherical aberration of this setup is exactly zero for all orders.
- Aplanatic means, that the linear coma vanishes and the imaging is free of coma for a small but finite field size. Show this property by using a small field of 2° for the current system. What is the largest present aberration?

Solution:

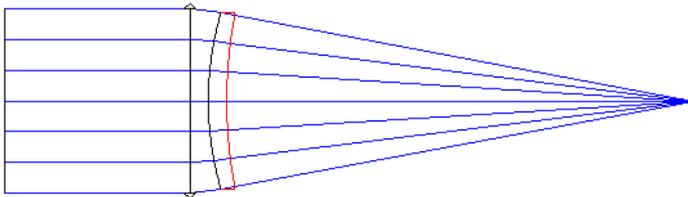
The initial focussing lens is established as follows:

	Comment	Radius	Thickness	Glass	Semi-Diameter	Conic	Par 0 (unused)	Par 1 (unused)
OBJ	Standard	Infinity	Infinity		0.0000000	0.0000000		
1	Standard	Infinity	10.0000000		5.0000000	0.0000000		
STO	Paraxial		0.0000000		5.0000000			50.0000000
3	Standard	Infinity	50.0000000		5.0000000	0.0000000		
IMA	Standard	Infinity	--		0.0000000	0.0000000		



- A lens with thickness 1 mm is placed 1 mm behind the lens. The first surface is made aplanatic by a solve, the second surface is mad concentric by choosing the solve

Surf#	Type	Comment	Radius	Thickness	Glass	Semi-Diameter	Conic	Par 0 (unused)	Par 1 (unused)
OBJ	Standard		Infinity	Infinity		0.0000000	0.0000000		
1	Standard		Infinity	10.0000000		5.0000000	0.0000000		
2	Paraxial		0.0000000			5.0000000			50.0000000
3	Standard		Infinity	1.0000000		5.0000000	0.0000000		
4	Standard		17.0884803 A	1.0000000	SP37	4.8303111	0.0000000		
5	Standard		25.2392874 N	25.2392874		4.6898661	0.0000000		
IMA	Standard		Infinity	-		1.776E-015	0.0000000		



If a single ray trace is performed, we get the direction cosine of the marginal ray to be 0.1858.

Surf	X-coordinate	Y-coordinate	Z-coordinate	X-cosine	Y-cosine
OBJ	Infinity	Infinity	Infinity	0.000000000	0.000000000
1	0.000000000E+000	5.000000000E+000	0.000000000E+000	0.000000000	0.000000000
2	0.000000000E+000	5.000000000E+000	0.000000000E+000	0.000000000	-0.0995037190
3	0.000000000E+000	5.000000000E+000	0.000000000E+000	0.000000000	-0.0995037190
4	0.000000000E+000	4.8303110655E+000	6.968894849E-001	0.000000000	-0.1858161068
5	0.000000000E+000	4.6898661252E+000	4.3983384906E-001	0.000000000	-0.1858161068
6	0.000000000E+000	1.7763568394E-015	0.000000000E+000	0.000000000	-0.1858161068

The best image position can be obtained by

1. Quick focus option
 2. Solve at the last surface with marginal ray height 0
 3. Pick up on the last (concentric) surface radius
 4. Optimizing the last thickness as a variable with minimal spot size
- b) If the Zernike polynomials are calculated, they are exactly zero for all orders.

2: Zernike Fringe Coefficients			
Update	Settings	Print	Window
Wavelength	:	0.5000	µm
Peak to Valley (to chief)	:	0.00000001	waves
Peak to Valley (to centroid)	:	0.00000001	waves
RMS (to chief)	:	0.00000000	waves
RMS (to centroid)	:	0.00000000	waves
Variance	:	0.00000000	waves
Strehl Ratio (Est)	:	1.00000000	
RMS fit error	:	0.00000000	waves
Maximum fit error	:	0.00000001	waves
Z 1	0.00000000	:	1
Z 2	0.00000000	:	(p) * COS (A)
Z 3	0.00000000	:	(p) * SIN (A)
Z 4	0.00000000	:	(2p ² - 1)
Z 5	0.00000000	:	(p ²) * COS (2A)
Z 6	0.00000000	:	(p ²) * SIN (2A)
Z 7	0.00000000	:	(3p ² - 2) p * COS (A)
Z 8	0.00000000	:	(3p ² - 2) p * SIN (A)
Z 9	0.00000000	:	(6p ⁴ - 6p ² + 1)
Z 10	0.00000000	:	(p ³) * COS (3A)
Z 11	0.00000000	:	(p ³) * SIN (3A)
Z 12	0.00000000	:	(4p ² -3) p ² * COS (2A)
Z 13	0.00000000	:	(4p ² -3) p ² * SIN (2A)
Z 14	0.00000000	:	(10p ⁴ - 12p ² + 3) p * COS
Z 15	0.00000000	:	(10p ⁴ - 12p ² + 3) p * SIN
Z 16	0.00000000	:	(20p ⁶ - 30p ⁴ + 12p ² - 1)
Z 17	0.00000000	:	(p ⁴) * COS (4A)
Z 18	0.00000000	:	(p ⁴) * SIN (4A)
Z 19	0.00000000	:	(5p ² - 4) p ³ * COS (3A)
Z 20	0.00000000	:	(5p ² - 4) p ³ * SIN (3A)
Z 21	0.00000000	:	(15p ⁴ - 20p ² + 6) p ² * C
Z 22	0.00000000	:	(15p ⁴ - 20p ² + 6) p ² * S
Z 23	0.00000000	:	(35p ⁶ - 60p ⁴ + 30p ² - 4)
Z 24	0.00000000	:	(35p ⁶ - 60p ⁴ + 30p ² - 4)
Z 25	0.00000000	:	(70p ⁸ - 140p ⁶ + 90p ⁴ - 20
Z 26	0.00000000	:	(p ⁵) * COS (5A)
Z 27	0.00000000	:	(p ⁵) * SIN (5A)
Z 28	0.00000000	:	(6p ² - 5) p ⁴ * COS (4A)
Z 29	0.00000000	:	(6p ² - 5) p ⁴ * SIN (4A)
Z 30	0.00000000	:	(21p ⁴ - 30p ² + 10) p ³ * C
Z 31	0.00000000	:	(21p ⁴ - 30p ² + 10) p ³ * S
Z 32	0.00000000	:	(56p ⁶ - 105p ⁴ + 60p ² - 10
Z 33	0.00000000	:	(56p ⁶ - 105p ⁴ + 60p ² - 10
Z 34	0.00000000	:	(126 p ⁸ - 280p ⁶ + 210p ⁴ -
Z 35	0.00000000	:	(126 p ⁸ - 280p ⁶ + 210p ⁴ -
Z 36	0.00000000	:	(252p ¹⁰ - 630p ⁸ + 560p ⁶ -

c) If a field of 2° is introduced and the Zernike coefficients are calculated for the field point in the image and behind the 4th surface (the aplanatic), we get the following picture:

Surface	:	4	
Field	:	2.0000 (deg)	
Wavelength	:	0.5000 µm	
Peak to Valley (to chief)	:	0.42436528 waves	
Peak to Valley (to centroid)	:	0.41948035 waves	
RMS (to chief)	:	0.11961177 waves	
RMS (to centroid)	:	0.11958708 waves	
Variance	:	0.01490107 waves squared	
Strehl Ratio (Est)	:	0.56859687	
RMS fit error	:	0.00000000 waves	
Maximum fit error	:	0.00000001 waves	
Z 1	0.20794872	:	1
Z 2	0.00000000	:	(p) * COS (A)
Z 3	0.00490450	:	(p) * SIN (A)
Z 4	0.20837470	:	(2p ² - 1)
Z 5	-0.00021636	:	(p ²) * COS (2A)
Z 6	0.00000000	:	(p ²) * SIN (2A)
Z 7	0.00000000	:	(3p ² - 2) p * COS (A)
Z 8	0.00246706	:	(3p ² - 2) p * SIN (A)
Z 9	0.00042809	:	(6p ⁴ - 6p ² + 1)
Z 10	0.00000000	:	(p ³) * COS (3A)
Z 11	-0.00000102	:	(p ³) * SIN (3A)
Z 12	-0.00001191	:	(4p ² -3) p ² * COS (2A)
Z 13	0.00000000	:	(4p ² -3) p ² * SIN (2A)
Z 14	0.00000000	:	(10p ⁴ - 12p ² + 3) p * COS (A)
Z 15	0.00000997	:	(10p ⁴ - 12p ² + 3) p * SIN (A)
Z 16	0.00000211	:	(20p ⁶ - 30p ⁴ + 12p ² - 1)
Z 17	0.00000000	:	(p ⁴) * COS (4A)
Z 18	0.00000000	:	(p ⁴) * SIN (4A)
Z 19	0.00000000	:	(5p ² - 4) p ³ * COS (3A)
Z 20	0.00000000	:	(5p ² - 4) p ³ * SIN (3A)

Surface	:	Image	
Field	:	2.0000 (deg)	
Wavelength	:	0.5000 µm	
Peak to Valley (to chief)	:	0.98906306 waves	
Peak to Valley (to centroid)	:	0.97834634 waves	
RMS (to chief)	:	0.22880861 waves	
RMS (to centroid)	:	0.22874623 waves	
Variance	:	0.05232484 waves squared	
Strehl Ratio (Est)	:	0.12672931	
RMS fit error	:	0.00000000 waves	
Maximum fit error	:	0.00000001 waves	
Z 1	0.34701665	:	1
Z 2	0.00000000	:	(p) * COS (A)
Z 3	0.01078681	:	(p) * SIN (A)
Z 4	0.34699634	:	(2p ² - 1)
Z 5	-0.27812765	:	(p ²) * COS (2A)
Z 6	0.00000000	:	(p ²) * SIN (2A)
Z 7	0.00000000	:	(3p ² - 2) p * COS (A)
Z 8	0.00539378	:	(3p ² - 2) p * SIN (A)
Z 9	-0.00001686	:	(6p ⁴ - 6p ² + 1)
Z 10	0.00000000	:	(p ³) * COS (3A)
Z 11	-0.00140829	:	(p ³) * SIN (3A)
Z 12	0.00065752	:	(4p ² -3) p ² * COS (2A)
Z 13	0.00000000	:	(4p ² -3) p ² * SIN (2A)
Z 14	0.00000000	:	(10p ⁴ - 12p ² + 3) p * COS (A)
Z 15	0.00000041	:	(10p ⁴ - 12p ² + 3) p * SIN (A)
Z 16	0.00000345	:	(20p ⁶ - 30p ⁴ + 12p ² - 1)
Z 17	0.00000683	:	(p ⁴) * COS (4A)
Z 18	0.00000000	:	(p ⁴) * SIN (4A)
Z 19	0.00000000	:	(5p ² - 4) p ³ * COS (3A)
Z 20	0.00000000	:	(5p ² - 4) p ³ * SIN (3A)

In the image, defocus (which is here in field the field curvature) and astigmatism are the dominating aberrations. Directly behind the aplanatic surface, only defocus has a considerable amount. This shows, that the concentric surface limits the system performance by astigmatism and field curvature. The change in the coma of the aplanatic surface is extremely small.

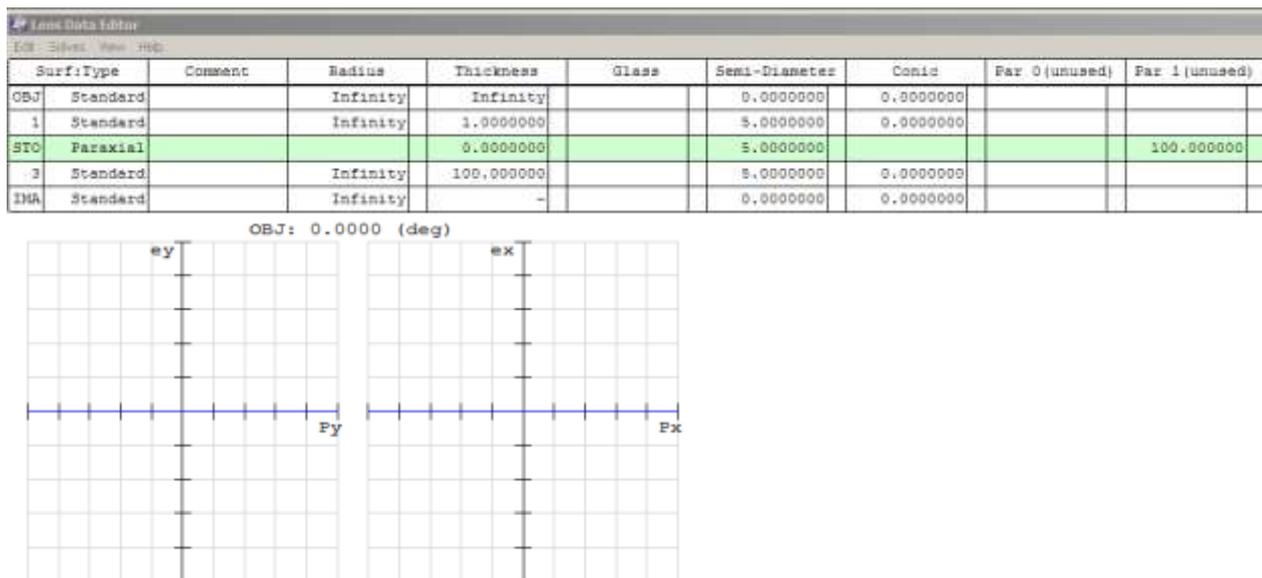
3.3 Transverse aberration curve

Establish a system with an ideal lens of focal length $f = 100$. mm, a wavelength of 600 nm and a collimated input bundle of diameter 10 mm.

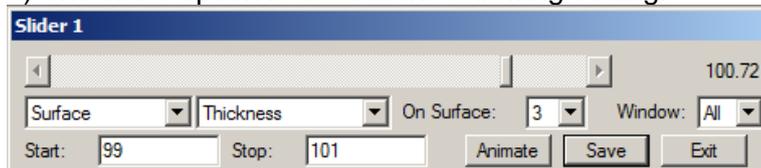
- Define the slider option with the image distance as a variable. Open the transverse aberration chart with a fixed scale of $100 \mu\text{m}$. What happens with the aberration curve, if the slider is moved ?
- Now introduce a thin glass plate in front of the perfect lens and define one side as a Zernike surface with coma as type (c8) of one wavelength. What is now the result for the moving slider ? Explain the result, if c7 is taken instead of c8.
- If now spherical aberration is introduced by the plate in selecting c9, what is seen for the defocus ? Open also a spot diagram with fixed scale. Determine the best image plane for an overall small spot diameter ?

Solution:

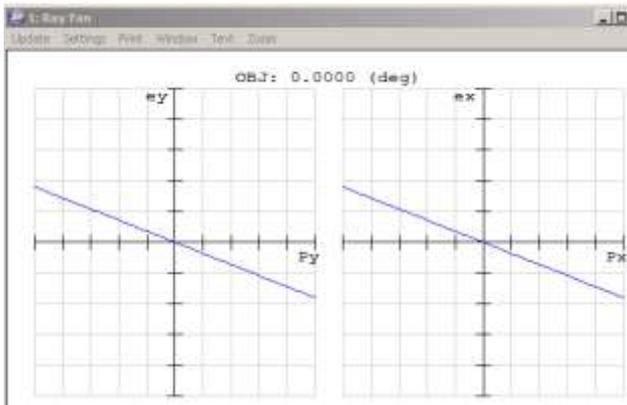
Setup of the initial system:



- The slider option is used in the following setting:



As a result, the transverse aberration curve is tilted corresponding to the interpretation of the transverse aberrations in the case of defocussing.



b) The system looks as follows:

Surf	Type	Comment	Radius	Thickness	Glass	Semi-Diameter	Conic	Par 0 (unused)	Par 1 (unused)	Par 2 (unused)	Par 3 (unused)	Par 4 (unused)
OBJ	Standard		Infinity	Infinity		0.000000	0.000000					
1	Standard		Infinity	0.000000		0.000000	0.000000					
2	Standard		Infinity	1.000000	SPK	0.000000	0.000000					
3	Zernike F..		Infinity	1.000000		0.000000	0.000000	1.000000	0			
STC	Paraxial		0.000000	0.000000		0.000000	0.000000		100.000000	1		
4	Standard		Infinity	100.000000		0.000000	0.000000					
DIA	Standard		Infinity	-		0.000000	0.000000					

Surf	Type	Maximal Term #	Bohm Radius	Zernike 1	Zernike 2	Zernike 3	Zernike 4	Zernike 5	Zernike 6	Zernike 7	Zernike 8
OBJ	Standard										
1	Standard										
2	Standard										
3	Zernike F..	9	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	1.000000
STC	Paraxial										

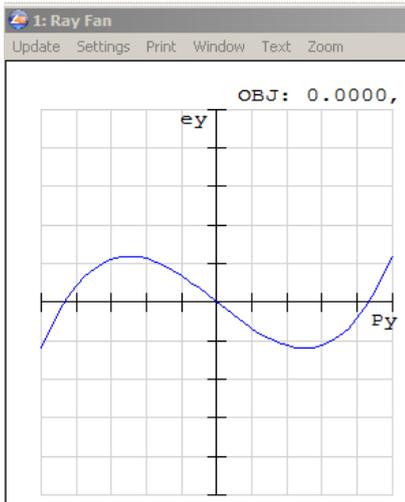
For a defocussing by the slider,



a strong parabolic shape due to the coma in the meridional plane is seen, which is tilted by the defocus.

If c7 is chosen the orientation of the coma is rotated and the parabola switches from the meridional to the sagittal plane.

c) For spherical aberration a cubic curve is superposed on the linear change. If the scale of the aberration curve is fixed to 300 μm , the best image plane is estimated to be at 101.44 mm. For this value, the residual zonal deviation has approximately the same size at the edge error.



3.4 Aspherical lens

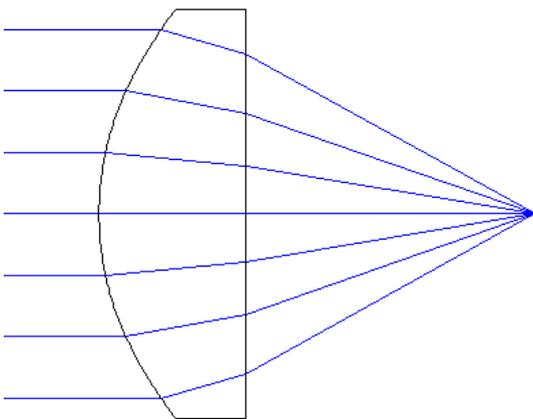
Load the catalog lens A20-18HPX from Asphericon.

- a) What is the residual correction in wave aberrations and in spot size ?
- b) Show, that the asphere is sensitive for finite field sizes by introducing a field with angle 0.1° . What is the dominating type of aberration ?
- c) What is the highest order of the chosen Taylor expansion ? Calculate the performance on axis , if the three highest expansion coefficients are set to zero. Is the system still diffraction limited ?

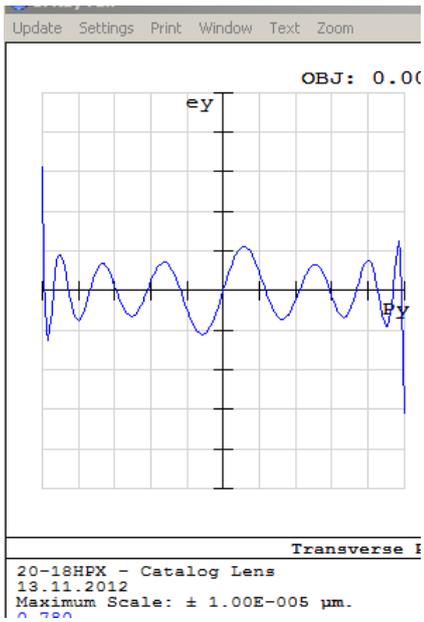
Solution:

The lens is established as follows:

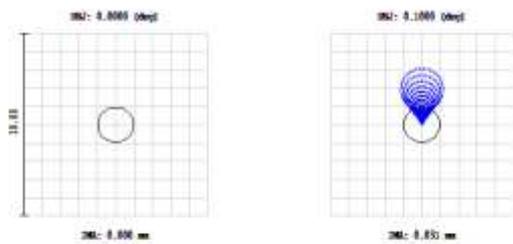
Part Type	Comment	Radius	Thickness	Class	Semi-Diameter	Conic	Par. 1 (umrad)	Par. 2 (umrad)	Par. 3 (umrad)	Par. 4 (umrad)	Par. 5 (umrad)	Par. 6 (umrad)	Par. 7 (umrad)
OBJ	Standard	Infinity	Infinity		0.000000	0.000000							
1	Standard	Infinity	1.000000		0.000000	0.000000							
* BETA Asp...	ASP-CRSP	13.980000	1.200000	S-LAKSH	10.200000	-1.000000	1.000000	3.488E-03	-2.32E-03	-1.04E-01	-8.17E-03	1.897E-01	1.389E-01
3*	Standard	Infinity	12.993424		0.000000	0.000000							
DIA	Standard	Infinity	-		0.22E-03	0.000000							



- a) The Zernike coefficients are zero in the available accuracy. The plots of the residual transverse aberrations shows a maximum value of 1 nm.



b) For the desired field, the spot shows coma in the range of the diffraction limit.



Surface: OBJ		Spot Diagram		
20-18HPX - Catalog Lens		Airy Radius: 0.9748 μm		
13.11.2012 Units are μm.				
Field :	1	2		
RMS radius :	0.000	1.187		
GEO radius :	0.000	3.134		Aspherical lens

The Zernike expansion is dominated by coma.

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Field                : 0.1000 (deg)
Wavelength           : 0.7800 µm
Peak to Valley (to chief) : 0.96880965 waves
Peak to Valley (to centroid) : 0.38180354 waves
RMS (to chief)       : 0.15898929 waves
RMS (to centroid)    : 0.06135497 waves
Variance             : 0.00376443 waves squared
Strehl Ratio (Est)   : 0.86190189

RMS fit error        : 0.00000023 waves
Maximum fit error    : 0.00000112 waves

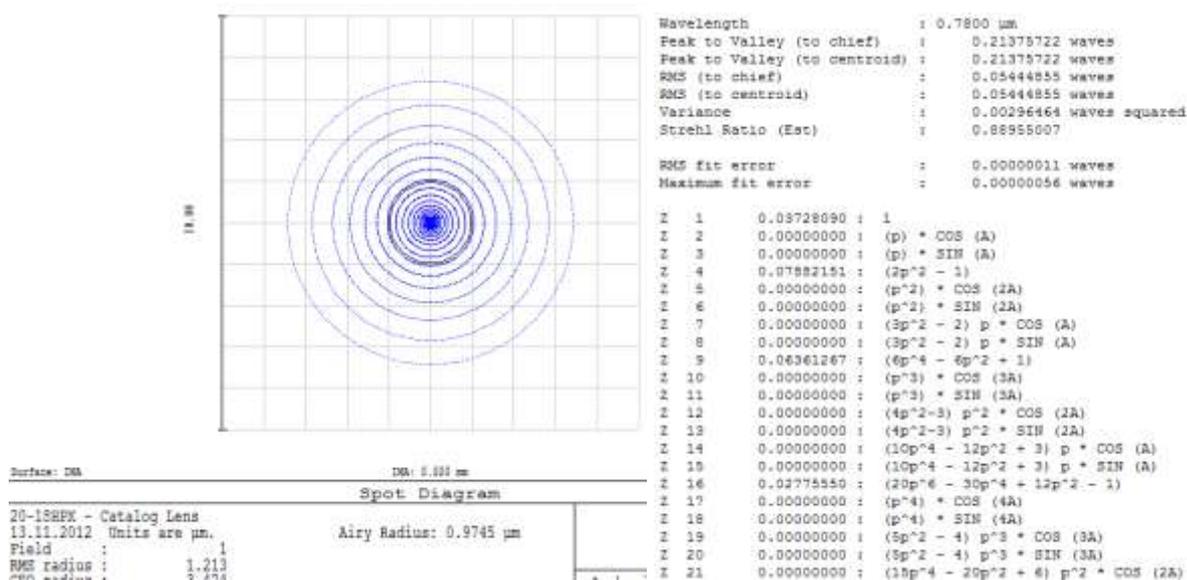
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Z 1 0.00496625 : 1
Z 2 0.00000000 : (p) * COS (A)
Z 3 -0.29390723 : (p) * SIN (A)
Z 4 0.00493362 : (2p^2 - 1)
Z 5 -0.00353074 : (p^2) * COS (2A)
Z 6 0.00000000 : (p^2) * SIN (2A)
Z 7 0.00000000 : (3p^2 - 2) p * COS (A)
Z 8 -0.17306735 : (3p^2 - 2) p * SIN (A)
Z 9 -0.00005601 : (6p^4 - 6p^2 + 1)
Z 10 0.00000000 : (p^3) * COS (3A)
Z 11 -0.00000114 : (p^3) * SIN (3A)
Z 12 0.00011117 : (4p^2-3) p^2 * COS (2A)
Z 13 0.00000000 : (4p^2-3) p^2 * SIN (2A)
Z 14 0.00000000 : (10p^4 - 12p^2 + 3) p * COS (A)
Z 15 -0.01747857 : (10p^4 - 12p^2 + 3) p * SIN (A)
Z 16 -0.00002371 : (20p^6 - 30p^4 + 12p^2 - 1)
Z 17 0.00000000 : (p^4) * COS (4A)
Z 18 0.00000000 : (p^4) * SIN (4A)
Z 19 0.00000000 : (5p^2 - 4) p^3 * COS (3A)
Z 20 0.00000007 : (5p^2 - 4) p^3 * SIN (3A)

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c) If the 10th to the 14th order coefficients are set to zero the following data are obtained:



With a Strehl number of 89% per definition the system is diffraction limited.